



High-frequency signals of oceans and atmosphere in Earth rotation

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TU Periodic polar motion – composition





- Ocean tides
 - Q1, O1, P1, K1,
 N2, M2, S2, K2
- Atmosphere tides
 - S1, S2
- Effect of lunisolar torque on triaxial Earth
 - PM: O1, P1, K1
 - UT1: M2, S2







 χ (motion)

 $\Delta UT1(\sigma)$

8 major tides

<u>ff</u>f

motion

 $u(\sigma), v(\sigma)$

8 major tides

Ocean tidal Earth rotation variations major terms



TU Atmospheric excitation



- Earth rotation excitation at daily and subdaily periods, investigated with different sets of AAM functions (from different ECMWF data classes)
 - Polar motion: amplitudes of S1 and S2 \sim 1 μ as
 - UT1: amplitudes of S1 and S2 < 0.5 μs



Pressure and wind terms partly balance each other, total net effect small

- Models of S1 and S2 are strongly dependent on the atmospheric model and the considered data time interval
- Amplitudes are small: pressure and wind effects, decisive for Earth rotation variations counterbalance each other.
- Poster XL121: Schindelegger et al.

High-frequency Earth rotation from VLBI

- Option 1: from highly resolved (1-2 h) ERP time series
 - 1st step: estimate celestial pole offsets (cpo) for all sessions
 - 2nd step: re-introduce cpo, fix nutation, estimate ERP with hourly resolution
 - 3rd step: remove low-frequency signal from the ERP time series, estimate amplitudes of selected periods in a least squares adjustment

ERP
time
series
$$\delta x_{p}(t) = \sum_{i=1}^{77} (-A_{i}^{+} - A_{i}^{-}) \cos(\xi_{i}(t)) + (B_{i}^{+} - B_{i}^{-}) \sin(\xi_{i}(t))$$
Tidal
coefficients
$$\delta y_{p}(t) = \sum_{i=1}^{77} (B_{i}^{+} + B_{i}^{-}) \cos(\xi_{i}(t)) + (A_{i}^{+} - A_{i}^{-}) \sin(\xi_{i}(t))$$

$$\delta UT1(t) = \sum_{i=1}^{77} U_{i}^{c} \cos(\xi_{i}(t)) + U_{i}^{s} \sin(\xi_{i}(t))$$

$$a_{ij} \dots 6 \text{ integer multipliers for each tide } i$$

$$a_{j} \dots 5 \text{ Delaunay variables } l, l', F, D, \Omega + (\text{GMST}+\pi)$$
as used in the IERS Conventions

High-frequency Earth rotation from VLBI

- Option 2: from demodulated ERP (complex demodulation technique, Herring & Dong (1994), Brzezinski (2012))
 - celestial pole offsets (nutation) can be estimated!
 - alternative ERP parameterisation:

$$\begin{bmatrix} x_p(t) \\ y_p(t) \end{bmatrix} = \sum_{\substack{k=-N \\ k\neq -1}}^{N} \begin{bmatrix} x_k(t) \\ y_k(t) \end{bmatrix} \cos(k\phi(t)) + \begin{bmatrix} y_k(t) \\ -x_k(t) \end{bmatrix} \sin(k\phi(t))$$
$$UT1(t) = \sum_{k=0}^{N} \frac{u_k^c(t)}{u_k^c(t)} \cos(k\phi(t)) + \frac{u_k^s(t)}{u_k^s(t)} \sin(k\phi(t)) \qquad \phi = GMST + \pi$$
$$k = 0 \Rightarrow \begin{bmatrix} x_0(t) \\ y_0(t) \end{bmatrix} \cos(0) + \begin{bmatrix} y_0(t) \\ x_0(t) \end{bmatrix} \sin(0)$$
$$\lim_{k \to \infty} \frac{u_k^c(t)}{u_k^c(t)} \cos(0) + \lim_{k \to \infty} \frac{u_k^c(t)}{u_k^c(t)} \sin(0)$$

k = 1...4 (diurnal, semidiurnal, terdiurnal, quarterdiurnal band)

High-frequency Earth rotation from VLBI

- Option 2: from demodulated ERP (complex demodulation technique)
 - preserves the amplitudes of periodic signals but shifts the frequencies
 - E.g. diurnal frequency band:



- Option 3: within a global solution
 - Accumulate single session normal equations
 - Estimate tidal terms directly together with other parameters such as station positions
- 3 sets of tidal amplitudes
 - Vienna VLBI Software VieVS: VLBI data 1984-2010
 - Polar and spin libration considered a priori
 - 40+3 diurnal, 30+3 semidiurnal, 1 terdiurnal (M3) tide/s

Option 1: highly resolved ERP time series Option 2: demodulated ERP time series Option 3: global solution



TU Polar motion w.r.t. IERS2010



TU Universal time w.r.t. IERS2010



RMS differences in terms of coefficients*

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ocean tide models		Model	IERS2010	TPXO7.2	HAM11a	VieVS 3	Artz et al.	ERP
		IERS2010		4.05	3.47	4.60	4.90	po mc
		TPXO7.2	0.32		3.07	4.21	4.53	lar otion [µas
VLBI global solution		HAM11a	0.45	0.30		5.31	5.86	
		VieVS 3	0.38	0.35	0.41		2.31	
	, ,	Artz et al.	0.38	0.35	0.43	0.16		
		ERP	universal time [µs]					

empirical model from a VLBI global solution, derived with Calc/Solve Software: Artz et al. (2011), *J Geod*.

* calculated only from the terms which are part of this model

- VLBI estimates show better mutual agreement than ocean tide models
- ERP variations based on TPXO7.2 fit best to the VLBI values, for polar motion as well as for UT1



- More recent ocean tide models do not necessarily agree better to empirical tidal ERP terms.
- There is room for improvement on the part of the modeling procedure?
- Extension with S1 tide from ocean tide models could help.
- S1 and S2 from atmospheric models should be interpreted with caution as they are strongly dependent on the used model and the considered data time interval.
- Combined analysis of VLBI and ring laser data
 - Poster XL 119, Nilsson et al.

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